

The HLLC Riemann Solver

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Abstract:

This lecture is about a method to solve approximately the Riemann problem for the Euler equations in order to derive a numerical flux for a conservative method:

The HLLC Riemann solver

REFERENCES:

E F Toro, M Spruce and W Speares.

Restoration of the contact surface in the HLL Riemann solver. Technical report CoA 9204. Department of Aerospace Science, College of Aeronautics, Cranfield Institute of Technology. UK. June, 1992.

E F Toro, M Spruce and W Speares.

Restoration of the contact surface in the Harten-Lax-van Leer Riemann solver. Shock Waves. Vol. 4, pages 25-34, 1994.

Consider the general Initial Boundary Value Problem (IBVP)

$$\left. \begin{array}{l} \text{PDEs} : \mathbf{U}_t + \mathbf{F}(\mathbf{U})_x = \mathbf{0}, \quad 0 \leq x \leq L, t > 0, \\ \text{ICs} : \mathbf{U}(x, 0) = \mathbf{U}^{(0)}(x), \\ \text{BCs} : \mathbf{U}(0, t) = \mathbf{U}_l(t), \quad \mathbf{U}(L, t) = \mathbf{U}_r(t), \end{array} \right\} \quad (1)$$

with appropriate boundary conditions, as solved by the explicit conservative scheme

$$\mathbf{U}_i^{n+1} = \mathbf{U}_i^n - \frac{\Delta t}{\Delta x} [\mathbf{F}_{i+\frac{1}{2}} - \mathbf{F}_{i-\frac{1}{2}}]. \quad (2)$$

The choice of numerical flux $\mathbf{F}_{i+\frac{1}{2}}$ determines the scheme. There are two classes of fluxes:

- ▶ Upwind or Godunov-type fluxes (wave propagation information used explicitly) and
- ▶ Centred or non-upwind (wave propagation information NOT used explicitly).

Godunov's flux (Godunov 1959) is

$$\mathbf{F}_{i+\frac{1}{2}} = \mathbf{F}(\mathbf{U}_{i+\frac{1}{2}}(0)) , \quad (3)$$

in which $\mathbf{U}_{i+\frac{1}{2}}(0)$ is the exact similarity solution $\mathbf{U}_{i+\frac{1}{2}}(x/t)$ of the Riemann problem

$$\left. \begin{aligned} \mathbf{U}_t + \mathbf{F}(\mathbf{U})_x &= \mathbf{0} , \\ \mathbf{U}(x, 0) &= \begin{cases} \mathbf{U}_L & \text{if } x < 0 , \\ \mathbf{U}_R & \text{if } x > 0 , \end{cases} \end{aligned} \right\} \quad (4)$$

evaluated at $x/t = 0$.

Example: 3D Euler equations.

$$\mathbf{U} = \begin{bmatrix} \rho \\ \rho u \\ \rho v \\ \rho w \\ E \end{bmatrix}, \quad \mathbf{F} = \begin{bmatrix} \rho u \\ \rho u^2 + p \\ \rho uv \\ \rho uw \\ u(E + p) \end{bmatrix}. \quad (5)$$

The piece-wise constant initial data, in terms of primitive variables, is

$$\mathbf{W}_L = \begin{bmatrix} \rho_L \\ u_L \\ v_L \\ w_L \\ p_L \end{bmatrix}, \quad \mathbf{W}_R = \begin{bmatrix} \rho_R \\ u_R \\ v_R \\ w_R \\ p_R \end{bmatrix}. \quad (6)$$

The Godunov flux $\mathbf{F}(\mathbf{U}_{i+\frac{1}{2}}(0))$ results from evaluation $\mathbf{U}_{i+\frac{1}{2}}(x/t)$ at $x/t = 0$, that is along the t -axis.

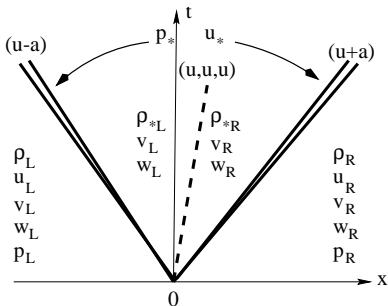


Fig. 1. Structure of the exact solution $\mathbf{U}_{i+\frac{1}{2}}(x/t)$ of the Riemann problem for the x -split three dimensional Euler equations. There are five wave families associated with the eigenvalues $u - a$, u (of multiplicity 3) and $u + a$.

Integral Relations

Consider the control volume $V = [x_L, x_R] \times [0, T]$ depicted in Fig. 2, with

$$x_L \leq TS_L, \quad x_R \geq TS_R, \quad (7)$$

S_L and S_R are the *fastest signal velocities* and T is a chosen time. The integral form of the conservation laws in (4) in V reads

$$\int_{x_L}^{x_R} \mathbf{U}(x, T) dx = \int_{x_L}^{x_R} \mathbf{U}(x, 0) dx + \int_0^T \mathbf{F}(\mathbf{U}(x_L, t)) dt - \int_0^T \mathbf{F}(\mathbf{U}(x_R, t)) dt. \quad (8)$$

Evaluation of the right-hand side of this expression gives

$$\int_{x_L}^{x_R} \mathbf{U}(x, T) dx = x_R \mathbf{U}_R - x_L \mathbf{U}_L + T(\mathbf{F}_L - \mathbf{F}_R), \quad (9)$$

where $\mathbf{F}_L = \mathbf{F}(\mathbf{U}_L)$ and $\mathbf{F}_R = \mathbf{F}(\mathbf{U}_R)$.

We call (9) the *consistency condition*.

Now split left-hand side of (8) into three integrals, namely

$$\int_{x_L}^{x_R} \mathbf{U}(x, T) dx = \int_{x_L}^{TS_L} \mathbf{U}(x, T) dx + \int_{TS_L}^{TS_R} \mathbf{U}(x, T) dx + \int_{TS_R}^{x_R} \mathbf{U}(x, T) dx$$

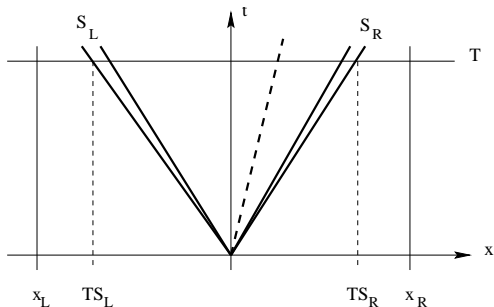


Fig. 2. Control volume $[x_L, x_R] \times [0, T]$ on $x-t$ plane. S_L and S_R are the fastest signal velocities arising from the solution of the Riemann problem.

Evaluate the first and third terms on the right-hand side to obtain

$$\int_{x_L}^{x_R} \mathbf{U}(x, T) dx = \int_{TS_L}^{TS_R} \mathbf{U}(x, T) dx + (TS_L - x_L) \mathbf{U}_L + (x_R - TS_R) \mathbf{U}_R. \quad (10)$$

Comparing (10) with (9) gives

$$\int_{TS_L}^{TS_R} \mathbf{U}(x, T) dx = T(S_R \mathbf{U}_R - S_L \mathbf{U}_L + \mathbf{F}_L - \mathbf{F}_R). \quad (11)$$

On division through by the length $T(S_R - S_L)$, which is the width of the wave system of the solution of the Riemann problem between the slowest and fastest signals at time T , we have

$$\frac{1}{T(S_R - S_L)} \int_{TS_L}^{TS_R} \mathbf{U}(x, T) dx = \frac{S_R \mathbf{U}_R - S_L \mathbf{U}_L + \mathbf{F}_L - \mathbf{F}_R}{S_R - S_L}. \quad (12)$$

Thus, the integral average of the exact solution of the Riemann problem between the slowest and fastest signals at time T is a known constant, provided that the signal speeds S_L and S_R are known; such constant is the right-hand side of (12) and we denote it by

$$\mathbf{U}^{hll} = \frac{S_R \mathbf{U}_R - S_L \mathbf{U}_L + F_L - F_R}{S_R - S_L}. \quad (13)$$

We now apply the integral form of the conservation laws to the left portion of Fig. 10.2, that is the control volume $[x_L, 0] \times [0, T]$. We obtain

$$\int_{TS_L}^0 \mathbf{U}(x, T) dx = -TS_L \mathbf{U}_L + T(\mathbf{F}_L - \mathbf{F}_{0L}), \quad (14)$$

where \mathbf{F}_{0L} is the flux $\mathbf{F}(\mathbf{U})$ along the t -axis. Solving for \mathbf{F}_{0L} we find

$$\mathbf{F}_{0L} = \mathbf{F}_L - S_L \mathbf{U}_L - \frac{1}{T} \int_{TS_L}^0 \mathbf{U}(x, T) dx. \quad (15)$$

Evaluation of the integral form of the conservation laws on the control volume $[0, x_R] \times [0, T]$ yields

$$\mathbf{F}_{0R} = \mathbf{F}_R - S_R \mathbf{U}_R + \frac{1}{T} \int_0^{TS_R} \mathbf{U}(x, T) dx . \quad (16)$$

The reader can easily verify that the equality

$$\mathbf{F}_{0L} = \mathbf{F}_{0R}$$

results in the *consistency condition* (9). All relations so far are exact, as we are assuming the exact solution of the Riemann problem.

The Harten-Lax-van Leer (HLL) Approximate Riemann Solver (1983).

$$\tilde{\mathbf{U}}(x, t) = \begin{cases} \mathbf{U}_L & \text{if } \frac{x}{t} \leq S_L, \\ \mathbf{U}^{hll} & \text{if } S_L \leq \frac{x}{t} \leq S_R, \\ \mathbf{U}_R & \text{if } \frac{x}{t} \geq S_R, \end{cases} \quad (17)$$

Fig. 3 shows the two-wave structure of this approximate Riemann solver.

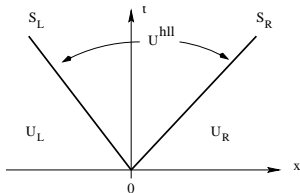


Fig. 3. Two-wave model. Approximate HLL Riemann solver. Solution in the *Star Region* consists of a single state \mathbf{U}^{hll} separated from data states by two waves of speeds S_L and S_R .

The HLL flux \mathbf{F}^{hll} for the subsonic case $S_L \leq 0 \leq S_R$ is found by inserting \mathbf{U}^{hll} in (13) into (15) or (16) to obtain

$$\mathbf{F}^{hll} = \mathbf{F}_L + S_L(\mathbf{U}^{hll} - \mathbf{U}_L), \quad (18)$$

or

$$\mathbf{F}^{hll} = \mathbf{F}_R + S_R(\mathbf{U}^{hll} - \mathbf{U}_R). \quad (19)$$

Use of (13) in (18) or (19) gives the HLL flux

$$\mathbf{F}^{hll} = \frac{S_R \mathbf{F}_L - S_L \mathbf{F}_R + S_L S_R (\mathbf{U}_R - \mathbf{U}_L)}{S_R - S_L} \quad (20)$$

for the subsonic case $S_L \leq 0 \leq S_R$.

The corresponding HLL intercell flux for the approximate Godunov method is then given by

$$\mathbf{F}_{i+\frac{1}{2}}^{hll} = \begin{cases} \mathbf{F}_L & \text{if } 0 \leq S_L, \\ \frac{S_R \mathbf{F}_L - S_L \mathbf{F}_R + S_L S_R (\mathbf{U}_R - \mathbf{U}_L)}{S_R - S_L}, & \text{if } S_L \leq 0 \leq S_R, \\ \mathbf{F}_R & \text{if } 0 \geq S_R. \end{cases} \quad (21)$$

- ▶ Given the speeds S_L and S_R we have an approximate intercell flux (21) to be used in the conservative formula (2) to produce an approximate Godunov method.
- ▶ A shortcoming of the HLL scheme, with its two-wave model, is exposed by contact discontinuities, shear waves and material interfaces, or any type of *intermediate waves*.

The HLLC Approximate Riemann Solver (Toro et al, 1992).

- ▶ The HLLC scheme is a modification of the HLL scheme whereby the missing contact and shear waves in the Euler equations are restored.
- ▶ HLLC for the Euler equations has a three-wave model

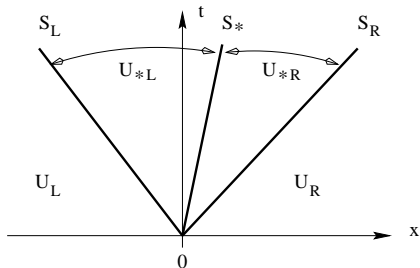


Fig. 4. HLLC approximate Riemann solver. Solution in the *Star Region* consists of two constant states separated from each other by a middle wave of speed S_* .

Useful Relations. Consider Fig. 2.

- ▶ Evaluation of the integral form of the conservation laws in the control volume reproduces the result of equation (12), even if variations of the integrand across the wave of speed S_* are allowed.
- ▶ Note that the consistency condition (9) effectively becomes the condition (12).
- ▶ By splitting the left-hand side of integral (12) into two terms we obtain

$$\frac{1}{T(S_R - S_L)} \int_{TS_L}^{TS_R} \mathbf{U}(x, T) dx = \mathbf{U}_{*L} + \mathbf{U}_{*R}, \quad (22)$$

where the following integral averages are introduced

$$\left. \begin{aligned} \mathbf{U}_{*L} &= \frac{1}{T(S_* - S_L)} \int_{TS_L}^{TS_*} \mathbf{U}(x, T) dx, \\ \mathbf{U}_{*R} &= \frac{1}{T(S_R - S_*)} \int_{TS_*}^{TS_R} \mathbf{U}(x, T) dx. \end{aligned} \right\} \quad (23)$$

Use of (23) into (22) and use of (12), make condition (9)

$$\left(\frac{S_* - S_L}{S_R - S_L} \right) \mathbf{U}_{*L} + \left(\frac{S_R - S_*}{S_R - S_L} \right) \mathbf{U}_{*R} = \mathbf{U}^{hll}, \quad (24)$$

The HLLC approximate Riemann solver is given as follows

$$\tilde{\mathbf{U}}(x, t) = \begin{cases} \mathbf{U}_L & , \text{ if } \frac{x}{t} \leq S_L, \\ \mathbf{U}_{*L} & , \text{ if } S_L \leq \frac{x}{t} \leq S_*, \\ \mathbf{U}_{*R} & , \text{ if } S_* \leq \frac{x}{t} \leq S_R, \\ \mathbf{U}_R & , \text{ if } \frac{x}{t} \geq S_R. \end{cases} \quad (25)$$

Now we seek a corresponding HLLC numerical flux of the form

$$\mathbf{F}_{i+\frac{1}{2}}^{hllc} = \begin{cases} \mathbf{F}_L & , \text{ if } 0 \leq S_L, \\ \mathbf{F}_{*L} & , \text{ if } S_L \leq 0 \leq S_*, \\ \mathbf{F}_{*R} & , \text{ if } S_* \leq 0 \leq S_R, \\ \mathbf{F}_R & , \text{ if } 0 \geq S_R, \end{cases} \quad (26)$$

with the intermediate fluxes \mathbf{F}_{*L} and \mathbf{F}_{*R} still to be determined, see Fig. 4. By integrating over appropriate control volumes we obtain

$$\mathbf{F}_{*L} = \mathbf{F}_L + S_L(\mathbf{U}_{*L} - \mathbf{U}_L), \quad (27)$$

$$\mathbf{F}_{*R} = \mathbf{F}_{*L} + S_*(\mathbf{U}_{*R} - \mathbf{U}_{*L}), \quad (28)$$

$$\mathbf{F}_{*R} = \mathbf{F}_R + S_R(\mathbf{U}_{*R} - \mathbf{U}_R). \quad (29)$$

These are three equations for the four unknowns vectors \mathbf{U}_{*L} , \mathbf{F}_{*L} , \mathbf{U}_{*R} , \mathbf{F}_{*R} .

We seek the solution for the two unknown intermediate fluxes \mathbf{F}_{*L} and \mathbf{F}_{*R} . There are more unknowns than equations and some extra conditions need to be imposed, in order to solve the algebraic problem. We impose

$$\left. \begin{aligned} p_{*L} &= p_{*R} = p_* , \\ u_{*L} &= u_{*R} = u_* , \end{aligned} \right\} \text{for pressure and normal velocity} \quad (30)$$

$$\left. \begin{aligned} v_{*L} &= v_L , & v_{*R} &= v_R , \\ w_{*L} &= w_L , & w_{*R} &= w_R . \end{aligned} \right\} \text{for tangential velocities} \quad (31)$$

Conditions (30), (31) are identically satisfied by the exact solution. In addition we impose

$$S_* = u_* \quad (32)$$

and thus if an estimate for S_* is known, the normal velocity component u_* in the *Star Region* is known.

Now equations (27) and (29) can be re-arranged as

$$S_L \mathbf{U}_{*L} - \mathbf{F}_{*L} = S_L \mathbf{U}_L - \mathbf{F}_L, \quad (33)$$

$$S_R \mathbf{U}_{*R} - \mathbf{F}_{*R} = S_R \mathbf{U}_R - \mathbf{F}_R, \quad (34)$$

where the right-hand sides of (33) and (34) are known constant vectors (data). We also note the useful relation

$$\mathbf{F}(\mathbf{U}) = u\mathbf{U} + p\mathbf{D}, \quad \mathbf{D} = [0, 1, 0, 0, u]^T. \quad (35)$$

Assuming S_L and S_R to be known and performing algebraic manipulations of the first and second components of equations (33)–(34) one obtains

$$p_{*L} = p_L + \rho_L(S_L - u_L)(S_{*} - u_L), \quad p_{*R} = p_R + \rho_R(S_R - u_R)(S_{*} - u_R). \quad (36)$$

From (30) $p_{*L} = p_{*R}$, which from (36) gives

$$S_* = \frac{p_R - p_L + \rho_L u_L (S_L - u_L) - \rho_R u_R (S_R - u_R)}{\rho_L (S_L - u_L) - \rho_R (S_R - u_R)}. \quad (37)$$

Manipulation of (33) and (34) and using p_{*L} and p_{*R} from (36) gives

$$\mathbf{F}_{*K} = \mathbf{F}_K + S_K (\mathbf{U}_{*K} - \mathbf{U}_K), \quad (38)$$

for $K=L$ and $K=R$, with the intermediate states given as

$$\mathbf{U}_{*K} = \rho_K \left(\frac{S_K - u_K}{S_K - S_*} \right) \left[\begin{array}{c} 1 \\ S_* \\ v_K \\ w_K \\ \frac{E_K}{\rho_K} + (S_* - u_K) \left[S_* + \frac{p_K}{\rho_K (S_K - u_K)} \right] \end{array} \right]. \quad (39)$$

The final choice of the HLLC flux is made according to (26).

Variation 1 of HLLC.

From equations (33) and (34) we may write the following solutions for the state vectors \mathbf{U}_{*L} and \mathbf{U}_{*R}

$$\mathbf{U}_{*K} = \frac{S_K \mathbf{U}_K - \mathbf{F}_K + p_{*K} \mathbf{D}_*}{S_L - S_*}, \quad \mathbf{D}_* = [0, 1, 0, 0, S_*], \quad (40)$$

with p_{*L} and p_{*R} as given by (36). Substitution of p_{*K} from (36) into (40) followed by use of (27) and (29) gives direct expressions for the intermediate fluxes as

$$\mathbf{F}_{*K} = \frac{S_*(S_K \mathbf{U}_K - \mathbf{F}_K) + S_K(p_K + \rho_L(S_K - u_K)(S_* - u_K))\mathbf{D}_*}{S_K - S_*}, \quad (41)$$

with the final choice of the HLLC flux made again according to (26).

Variation 2 of HLLC.

A different HLLC flux is obtained by assuming a single mean pressure value in the *Star Region*, and given by the arithmetic average of the pressures in (36), namely

$$P_{LR} = \frac{1}{2}[\rho_L + \rho_R + \rho_L(S_L - u_L)(S_* - u_L) + \rho_R(S_R - u_R)(S_* - u_R)] . \quad (42)$$

Then the intermediate state vectors are given by

$$\mathbf{U}_{*K} = \frac{S_K \mathbf{U}_K - \mathbf{F}_K + P_{LR} \mathbf{D}_*}{S_K - S_*} . \quad (43)$$

Substitution of these into (27) and (29) gives the fluxes \mathbf{F}_{*L} and \mathbf{F}_{*R} as

$$\mathbf{F}_{*K} = \frac{S_*(S_K \mathbf{U}_K - \mathbf{F}_K) + S_K P_{LR} \mathbf{D}_*}{S_K - S_*} . \quad (44)$$

Again the final choice of HLLC flux is made according to (26).

Remarks.

- ▶ The original HLLC formulation (38)–(39) enforces the condition $p_{*L} = p_{*R}$, which is satisfied by the exact solution.
- ▶ In the alternative HLLC formulation (41) we relax such condition, being more consistent with the pressure approximations (36).
- ▶ There is limited practical experience with the alternative HLLC formulations (41) and (44).
- ▶ *General equation of state.* All manipulations, assuming that wave speed estimates for S_L and S_R are available, are valid for any equation of state; this only enters when prescribing estimates for S_L and S_R .

Multidimensional multicomponent flow.

Consider the advection of a *chemical species* of concentrations q_l by the normal flow speed u . Then we can write the following advection equation

$$\partial_t q_l + u \partial_x q_l = 0, \text{ for } l = 1, \dots, m.$$

Note that these equations are written in non-conservative form. However, by combining these with the continuity equation we obtain a conservative form of these equations, namely

$$(\rho q_l)_t + (\rho u q_l)_x = 0, \text{ for } l = 1, \dots, m.$$

The eigenvalues of the enlarged system are as before, with the exception of $\lambda_2 = u$, which now, in three space dimensions, has multiplicity $m + 3$.

These conservation equations can then be added as new components to the conservation equations in (1) or (4), with the enlarged vectors of conserved variables and fluxes given as

$$\mathbf{U} = \begin{bmatrix} \rho \\ \rho u \\ \rho v \\ \rho w \\ E \\ \rho q_1 \\ \dots \\ \rho q_l \\ \dots \\ \rho q_m \end{bmatrix}, \quad \mathbf{F} = \begin{bmatrix} \rho u \\ \rho u^2 + p \\ \rho uv \\ \rho uw \\ u(E + p) \\ \rho u q_1 \\ \dots \\ \rho u q_l \\ \dots \\ \rho u q_m \end{bmatrix}. \quad (45)$$

The HLLC flux accommodates these new equations in a very natural way, and nothing special needs to be done. If the HLLC flux (38) is used, with \mathbf{F} as in (45), then the intermediate state vectors are given by

$$\mathbf{u}_{*K} = \rho_K \left(\frac{S_K - u_K}{S_K - S_*} \right) \left[\begin{array}{c} \frac{E_K}{\rho_K} + (S_* - u_K) \left[S_* + \frac{p_K}{\rho_K(S_K - u_K)} \right] \\ v_K \\ w_K \\ (q_1)_K \\ \dots \\ (q_l)_K \\ \dots \\ (q_m)_K \end{array} \right]. \quad (46)$$

for $K = L$ and $K = R$.

Wave Speed Estimates

We need estimates S_L , S_* and S_R . Davis (1988) suggested

$$S_L = u_L - a_L, \quad S_R = u_R + a_R, \quad (47)$$

$$S_L = \min \{u_L - a_L, u_R - a_R\}, \quad S_R = \max \{u_L + a_L, u_R + a_R\}. \quad (48)$$

Both Davis (1988) and Einfeldt (1988), proposed

$$S_L = \tilde{u} - \tilde{a}, \quad S_R = \tilde{u} + \tilde{a}, \quad (49)$$

\tilde{u} and \tilde{a} are the Roe-average particle and sound speeds respectively

$$\tilde{u} = \frac{\sqrt{\rho_L}u_L + \sqrt{\rho_R}u_R}{\sqrt{\rho_L} + \sqrt{\rho_R}}, \quad \tilde{a} = \left[(\gamma - 1) \left(\tilde{H} - \frac{1}{2}\tilde{u}^2 \right) \right]^{1/2}, \quad (50)$$

with the enthalpy $H = (E + p)/\rho$ approximated as

$$\tilde{H} = \frac{\sqrt{\rho_L}H_L + \sqrt{\rho_R}H_R}{\sqrt{\rho_L} + \sqrt{\rho_R}}. \quad (51)$$

Einfeldt (1988) proposed the estimates

$$S_L = \bar{u} - \bar{d}, \quad S_R = \bar{u} + \bar{d}, \quad (52)$$

for his HLLE solver, where

$$\bar{d}^2 = \frac{\sqrt{\rho_L} a_L^2 + \sqrt{\rho_R} a_R^2}{\sqrt{\rho_L} + \sqrt{\rho_R}} + \eta_2 (u_R - u_L)^2 \quad (53)$$

and

$$\eta_2 = \frac{1}{2} \frac{\sqrt{\rho_L} \sqrt{\rho_R}}{(\sqrt{\rho_L} + \sqrt{\rho_R})^2}. \quad (54)$$

These wave speed estimates are reported to lead to effective and robust Godunov-type schemes.

One-wave model.

Consider a **one-wave model** with single speed $S^+ > 0$.

- ▶ **Rusanov:** By choosing $S_L = -S^+$ and $S_R = S^+$ in the HLL flux (20) one obtains a Rusanov flux (1961)

$$\mathbf{F}_{i+1/2} = \frac{1}{2}(\mathbf{F}_L + \mathbf{F}_R) - \frac{1}{2}S^+(\mathbf{U}_R - \mathbf{U}_L). \quad (55)$$

- ▶ **Lax-Friedrichs:** Another possibility is $S^+ = S_{max}^n$, the wave speed for imposing the CFL condition, which satisfies

$$S_{max}^n = \frac{C_{cfl}\Delta x}{\Delta t}, \quad (56)$$

where C_{cfl} is the CFL coefficient. For $C_{cfl} = 1$, $S^+ = \frac{\Delta x}{\Delta t}$, which gives the Lax-Friedrichs numerical flux

$$\mathbf{F}_{i+1/2} = \frac{1}{2}(\mathbf{F}_L + \mathbf{F}_R) - \frac{1}{2}\frac{\Delta x}{\Delta t}(\mathbf{U}_R - \mathbf{U}_L). \quad (57)$$

Pressure-Based Wave Speed Estimates

Toro et al. (1994) suggested to first find an estimate p_* for the pressure in the *Star Region* and then take

$$S_L = u_L - a_L q_L, \quad S_R = u_R + a_R q_R, \quad (58)$$

$$q_K = \begin{cases} 1 & \text{if } p_* \leq p_K \\ \left[1 + \frac{\gamma + 1}{2\gamma} (p_*/p_K - 1) \right]^{1/2} & \text{if } p_* > p_K. \end{cases} \quad (59)$$

- ▶ This choice discriminates between shocks and rarefactions.
- ▶ If the K wave is a rarefaction then the speed S_K is the speed of the head of the rarefaction, the fastest signal.
- ▶ If the K wave is a shock wave then the speed is an approximation of the shock speed.

A simple, acoustic type approximation for pressure is (Toro, 1991)

$$p_* = \max(0, p_{pvrs}), \quad p_{pvrs} = \frac{1}{2}(p_L + p_R) - \frac{1}{2}(u_R - u_L)\bar{\rho}\bar{a}, \quad (60)$$

where

$$\bar{\rho} = \frac{1}{2}(\rho_L + \rho_R), \quad \bar{a} = \frac{1}{2}(a_L + a_R). \quad (61)$$

Another choice is furnished by the Two-Rarefaction Riemann solver, namely

$$p_* = p_{tr} = \left[\frac{a_L + a_R - \frac{\gamma-1}{2}(u_R - u_L)}{a_L/p_L^z + a_R/p_R^z} \right]^{1/z}, \quad (62)$$

where

$$P_{LR} = \left(\frac{p_L}{p_R} \right)^z; \quad z = \frac{\gamma-1}{2\gamma}. \quad (63)$$

The Two-Shock Riemann solver gives

$$p_* = p_{ts} = \frac{g_L(p_0)p_L + g_R(p_0)p_R - \Delta u}{g_L(p_0) + g_R(p_0)}, \quad (64)$$

where

$$g_K(p) = \left[\frac{A_K}{p + B_K} \right]^{1/2}, \quad p_0 = \max(0, p_{pvrs}), \quad (65)$$

for $K = L$ and $K = R$.

Summary of HLLC Fluxes

- ▶ *Step I: pressure estimate p_* .*
- ▶ *Step II: wave speed estimates:*

$$S_L = u_L - a_L q_L, \quad S_R = u_R + a_R q_R, \quad (66)$$

with

$$q_K = \begin{cases} 1 & \text{if } p_* \leq p_K \\ \left[1 + \frac{\gamma + 1}{2\gamma} (p_*/p_K - 1) \right]^{1/2} & \text{if } p_* > p_K. \end{cases} \quad (67)$$

and

$$S_* = \frac{p_R - p_L + \rho_L u_L (S_L - u_L) - \rho_R u_R (S_R - u_R)}{\rho_L (S_L - u_L) - \rho_R (S_R - u_R)}. \quad (68)$$

- ▶ *Step III: HLLC flux.* Compute the HLLC flux, according to

$$\mathbf{F}_{i+\frac{1}{2}}^{hllc} = \begin{cases} \mathbf{F}_L & \text{if } 0 \leq S_L, \\ \mathbf{F}_{*L} & \text{if } S_L \leq 0 \leq S_*, \\ \mathbf{F}_{*R} & \text{if } S_* \leq 0 \leq S_R, \\ \mathbf{F}_R & \text{if } 0 \geq S_R, \end{cases} \quad (69)$$

$$\mathbf{F}_{*K} = \mathbf{F}_K + S_K(\mathbf{U}_{*K} - \mathbf{U}_K) \quad (70)$$

and

$$\mathbf{U}_{*K} = \rho_K \left(\frac{S_K - u_K}{S_K - S_*} \right) \begin{bmatrix} 1 \\ S_* \\ v_K \\ w_K \\ \frac{E_K}{\rho_K} + (S_* - u_K) \left[S_* + \frac{\rho_K}{\rho_K(S_K - u_K)} \right] \end{bmatrix} \cdot \quad (71)$$

There are two variants of the HLLC flux in the third step, as seen below.

- ▶ *Step III: HLLC flux, Variant 1.* Compute the numerical fluxes as

$$\mathbf{F}_{*K} = \frac{S_*(S_K \mathbf{U}_K - \mathbf{F}_K) + S_K(\rho_K + \rho_L(S_K - u_K)(S_* - u_K))\mathbf{D}_*}{S_K - S_*},$$

$$\mathbf{D}_* = [0, 1, 0, 0, S_*]^T,$$
(72)

and the final HLLC flux chosen according to (69).

- ▶ *Step III: HLLC flux, Variant 2.* Compute the numerical fluxes as

$$\mathbf{F}_{*K} = \frac{S_*(S_K \mathbf{U}_K - \mathbf{F}_K) + S_K P_{LR} \mathbf{D}_*}{S_K - S_*},$$
(73)

with \mathbf{D}_* as in (72) and

$$P_{LR} = \frac{1}{2} [p_L + p_R + \rho_L(S_L - u_L)(S_* - u_L) + \rho_R(S_R - u_R)(S_* - u_R)].$$
(74)

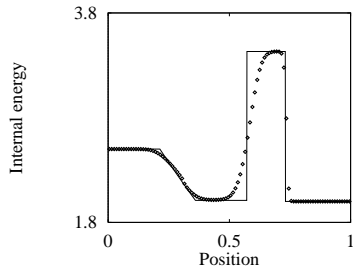
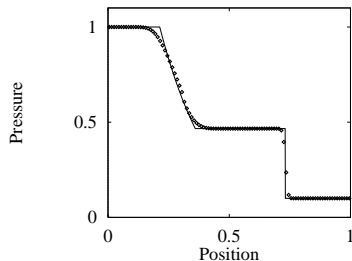
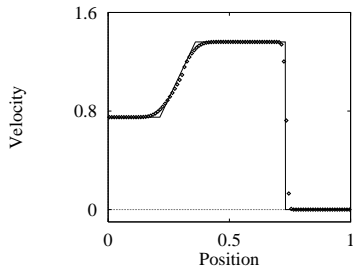
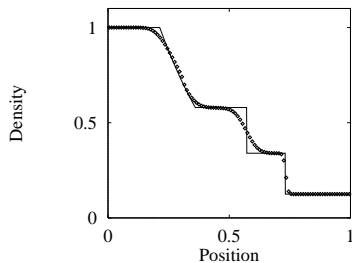
The final HLLC flux is chosen according to (69).

Numerical Results

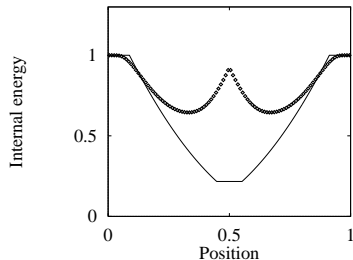
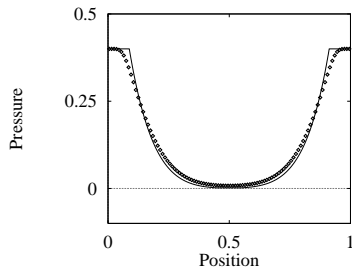
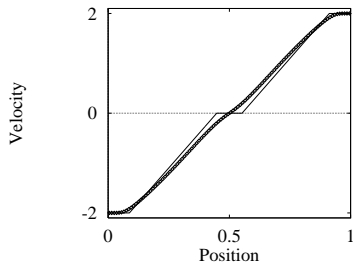
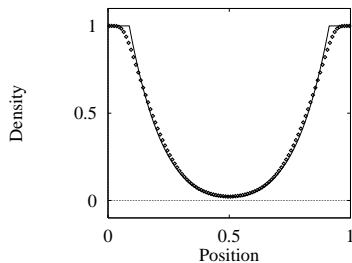
Test problems:

Test	ρ_L	u_L	p_L	ρ_R	u_R	p_R
1	1.0	0.75	1.0	0.125	0.0	0.1
2	1.0	-2.0	0.4	1.0	2.0	0.4
3	1.0	0.0	1000.0	1.0	0.0	0.01
4	5.99924	19.5975	460.894	5.99242	-6.19633	46.0950
5	1.0	-19.59745	1000.0	1.0	-19.59745	0.01
6	1.4	0.0	1.0	1.0	0.0	1.0
7	1.4	0.1	1.0	1.0	0.1	1.0

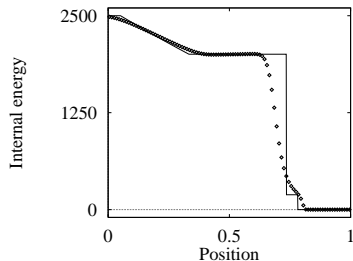
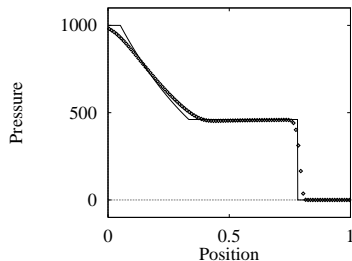
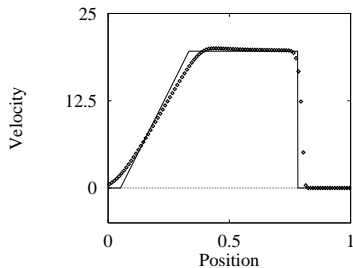
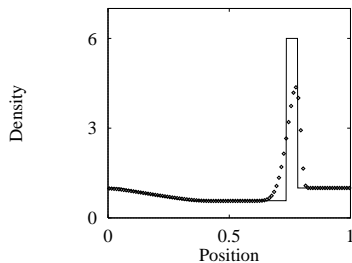
Table 1. Data for seven test problems with exact solution



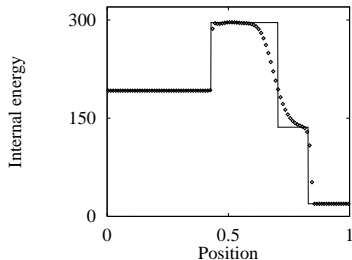
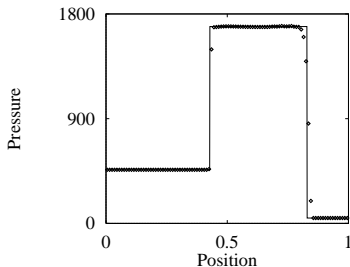
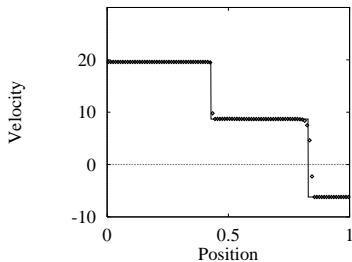
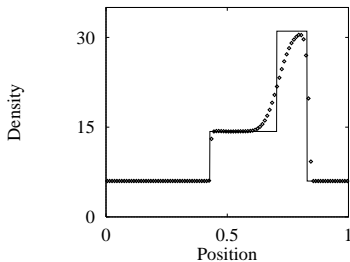
Godunov's method with HLLC Riemann solver applied to Test 1, with $x_0 = 0.3$. Numerical (symbol) and exact (line) solutions are compared at time 0.2.



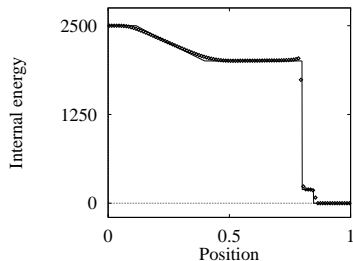
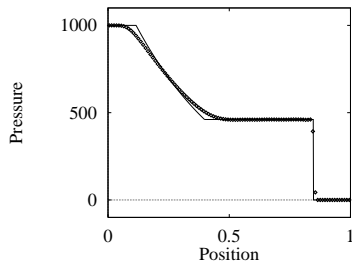
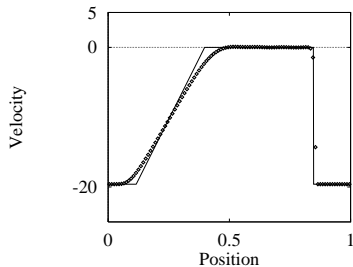
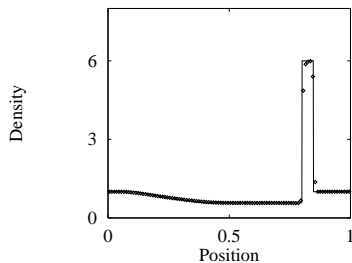
Godunov's method with HLLC Riemann solver applied to Test 2, with $x_0 = 0.5$. Numerical (symbol) and exact (line) solutions are compared at time 0.15.



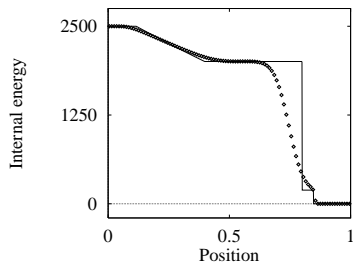
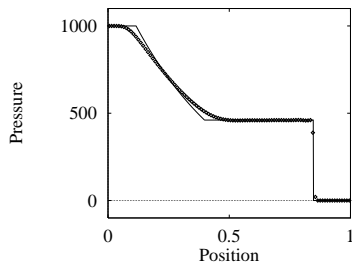
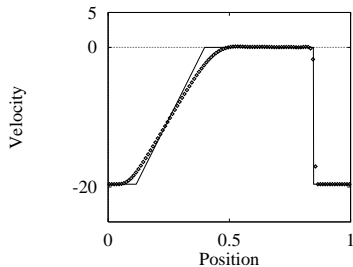
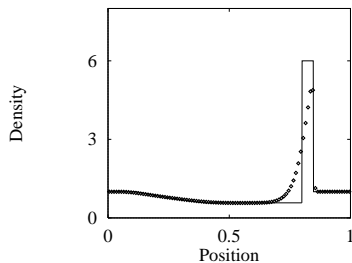
Godunov's method with HLLC Riemann solver applied to Test 3, with $x_0 = 0.5$. Numerical (symbol) and exact (line) solutions are compared at time 0.012.



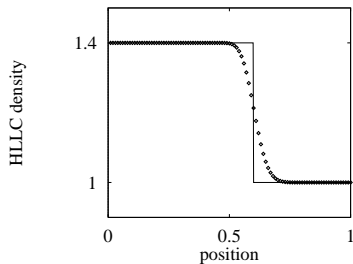
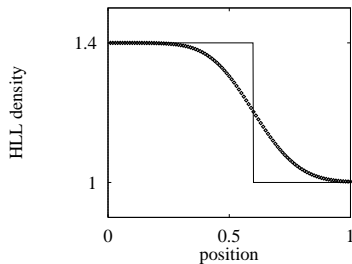
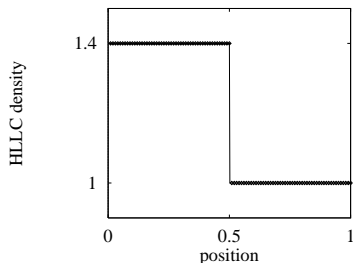
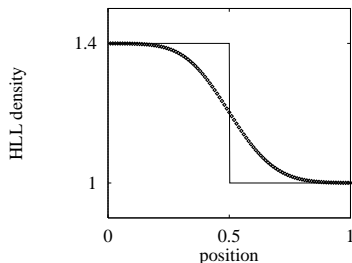
Godunov's method with HLLC Riemann solver applied to Test 4, with $x_0 = 0.4$. Numerical (symbol) and exact (line) solutions are compared at time 0.035.



Godunov's method with HLLC Riemann solver applied to Test 5, with $x_0 = 0.8$. Numerical (symbol) and exact (line) solutions are compared at time 0.012.



Godunov's method with HLL Riemann solver applied to Test 5, with $x_0 = 0.8$. Numerical (symbol) and exact (line) solutions are compared at time 0.012.



Godunov's method with HLL (left) and HLLC (right) Riemann solvers applied to Tests 6 and 7. Numerical (symbol) and exact (line) solutions are compared at time 2.0.

Closing Remarks:

- ▶ We have presented HLLC for the Euler equations.
- ▶ For the 2D shallow water equations see Toro E F Shock capturing methods for free-surface shallow flows. Wiley and Sons, 2001.
- ▶ For Turbulent flow applications (implicit version of HLLC), see Batten, Leschziner and Goldberg (1997).
- ▶ For extensions to MHD equations see Gurski (2004), Li (2005), Mignone et al. (2006++).
- ▶ For application to two-phase flow see Tokareva and Toro, JCP (2010).
- ▶ For extensions see Takahiro (2005) and Bouchut (2007), Mignone (2005).