

Waves and related processes in Geophysical Fluid Dynamics

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Non-linear Rossby waves in the QG RSW model on the β -plane

Non-dimensional equations of motion :

$$\nabla^2 \psi_t - \psi_t + \epsilon \mathcal{J}(\psi, \nabla^2 \psi) + \psi_x = 0,$$

ϵ - non-linearity parametre, $\epsilon \rightarrow 0$. Asymptotic expansion in non-linearity parametre. Solution - **asymtotic series** :

$$\psi = \psi^{(0)} + \epsilon \psi^{(1)} + \dots$$

Order zero : linear Rossby waves.

$$\nabla^2 \psi_t^{(0)} - \psi_t^{(0)} + \psi_x^{(0)} = 0, \Rightarrow$$

$$\psi^{(0)} = \sum_i A_i e^{i(\mathbf{k}_i \cdot \mathbf{x} - \omega(\mathbf{k}_i)t)} + \text{c.c.}, \quad \omega(\mathbf{k}) = -\frac{k}{k^2 + l^2 + 1}, \quad \mathbf{k} = (k, l). \quad (1)$$

Order one : first non-linear correction :

$$\nabla^2 \psi_t^{(1)} - \psi_t^{(1)} + \psi_x^{(1)} = -\mathcal{J}(\psi^{(0)}, \nabla^2 \psi^{(0)}), \quad (2)$$

Term in the r.h.s. :

$$\begin{aligned} & \sum_{i,j} A_i A_j \left[(k_i l_j - k_j l_i) \mathbf{k}_j^2 \right] e^{i[(\mathbf{k}_i + \mathbf{k}_j) \cdot \mathbf{x} - (\omega(\mathbf{k}_i) + \omega(\mathbf{k}_j))t]} \quad (3) \\ - & \sum_{i,j} A_i A_j^* \left[(k_i l_j - k_j l_i) \mathbf{k}_j^2 \right] e^{i[(\mathbf{k}_i - \mathbf{k}_j) \cdot \mathbf{x} - (\omega(\mathbf{k}_i) - \omega(\mathbf{k}_j))t]} + \text{c.c.} \end{aligned}$$

Integrability conditions : solution $\psi^{(1)}$ should be **bounded**

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Integrability conditions :

$$\forall \hat{\psi} : \nabla^2 \hat{\psi}_t - \hat{\psi}_t + \hat{\psi}_x = 0,$$

$$\int_{-\infty}^{\infty} dt \int_{-\infty}^{\infty} dx \int_{-\infty}^{\infty} dy \hat{\psi}^* \left(\nabla^2 \psi_t^{(1)} - \psi_t^{(1)} + \psi_x^{(1)} \right) = 0.$$

Therefore :

$$\int_{-\infty}^{\infty} dt \int_{-\infty}^{\infty} dx \int_{-\infty}^{\infty} dy \hat{\psi}^* \left(\mathcal{J}(\psi, \nabla^2 \psi) \right) = 0. \quad (4)$$

- orthogonality of the r.h.s. to the eigen-vectors of the zero-order linear operator.

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Necessarily : $\hat{\psi} \propto e^{i(\hat{\mathbf{k}} \cdot \mathbf{x} - \omega(\hat{\mathbf{k}})t)}$, and (4) becomes :

$$\int_{-\infty}^{\infty} dt dx dy e^{i[(\mathbf{k}_i + \mathbf{k}_j - \hat{\mathbf{k}}) \cdot \mathbf{x} - (\omega(\mathbf{k}_i) + \omega(\mathbf{k}_j) - \omega(\hat{\mathbf{k}}))t]} \sum_{i,j} A_i A_j [(k_{ilj} - k_{jlj}) \mathbf{k}_j^2] \cdot$$

$$\int_{-\infty}^{\infty} dt dx dy e^{i[(\mathbf{k}_i - \mathbf{k}_j - \hat{\mathbf{k}}) \cdot \mathbf{x} - (\omega(\mathbf{k}_i) - \omega(\mathbf{k}_j) - \omega(\hat{\mathbf{k}}))t]} \sum_{i,j} A_i A_j^* [(k_{ilj} - k_{jlj}) \mathbf{k}_j^2] \cdot + c.c. = 0$$

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Calculus of generalised functions :

$$\int_{-\infty}^{\infty} dx e^{ikx} = \delta(k) - \text{Dirac's delta-function.} \quad (5)$$

Generalisation of

$$\int_0^{2\pi} dx e^{ikx} = \delta_{k0} - \text{tensor delta of Kronecker} \quad (6)$$

for periodic boundary conditions.

Resonances :

Non-zero contributions :

$$\mathbf{k}_i \pm \mathbf{k}_j = \hat{\mathbf{k}}, \quad \omega(\mathbf{k}_i) \pm \omega(\mathbf{k}_j) = \omega(\hat{\mathbf{k}}). \quad (7)$$

three-wave **resonances**, **resonant triads**.

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Elimination of resonances

If $\exists \hat{\mathbf{k}}$ which verifies (7) the first non-linear correction is not bounded \Rightarrow asymptotic procedure is not self-consistent : resonances should be "killed".

Introducing slow evolution of the amplitudes :

$$\begin{aligned} \partial_t &\rightarrow \partial_t + \epsilon \partial_T \Rightarrow & (8) \\ \nabla^2 \psi_t^{(1)} - \psi_t^{(1)} + \psi_x^{(1)} &= -\nabla^2 \psi_T^{(0)} - \psi_T^{(0)} - \mathcal{J}(\psi^{(0)}, \nabla^2 \psi^{(0)}) \end{aligned}$$

New contribution in the r.h.s. :

$$\sum_i A_{iT} e^{i(\mathbf{k}_i \cdot \mathbf{x} - \omega(\mathbf{k}_i)t)} + c.c. \Rightarrow \quad (9)$$

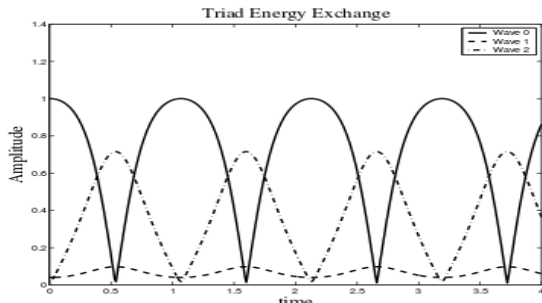
Possibility of **compensation** of resonant contributions by **slow evolution of amplitudes**.

A resonant triad :

$$\mathbf{k}_1 + \mathbf{k}_2 = \mathbf{k}_3, \quad \omega(\mathbf{k}_1) + \omega(\mathbf{k}_2) = \omega(\mathbf{k}_3), \quad (10)$$

$$\begin{aligned} \dot{A}_3 &= c(\mathbf{k}_1, \mathbf{k}_2) A_1 A_2, \\ \dot{A}_2 &= c(\mathbf{k}_3, -\mathbf{k}_1) A_3 A_1^*, \\ \dot{A}_1 &= c(\mathbf{k}_3, -\mathbf{k}_2) A_3 A_2^*, \end{aligned} \quad (11)$$

where $c(\mathbf{k}_1, \mathbf{k}_2) = \hat{\mathbf{z}} \cdot (\mathbf{k}_1 \wedge \mathbf{k}_2) \mathbf{k}_2^2$ - **interaction coefficients** .
This is an **integrable system** (in elliptic functions). Energy is conserved and redistributed among three waves.



Waveguides in RSW

Equator

- ▶ Trapped waves - baroclinic : Kelvin (nondispersive), IG, Rossby, Yanai (dispersive),
- ▶ Free waves : barotropic Rossby (dispersive)

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Coast/Escarpment/Ridge

- ▶ Trapped waves
 - ▶ Coast : shelf and edge waves (dispersive) and Kelvin (nondispersive for vertical edge),
 - ▶ Escarpment/Ridge : topographic Rossby(dispersive)
- ▶ Free waves (incident + reflected + transmitted) : IG (dispersive).

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The ideology

- ▶ Interactions between free waves and trapped waves (waveguide modes) may be **resonant**.
- ▶ If so the incoming free waves may **resonantly excite** waveguide modes.
- ▶ The resonant growth, if confirmed by "naive" asymptotic expansions, should be nonlinearly or dissipatively saturated in one way or another in the "improved" expansions leading to **coherent structures** formation.

Remark :

One of the waveguide modes forming a resonant triad may have infinite wavelength, i.e. be a **mean flow**.

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- ▶ Take a **waveguide** in (multi-layer) RSW
- ▶ Analyse the **dispersion relations** for free and trapped waves and check if **resonances** between them are possible (algebra may be tough)
- ▶ Apply straightforward ("naive") asymptotic expansions in nonlinearity parameter to the flow consisting of resonating waves and "eliminate" resonances. **Caveat** : *Normalizations are different for trapped and free waves - should be careful.*
- ▶ Check a possibility of **parametric resonance** in slow-time evolution equations and growth of amplitude(s) of the trapped wave(s).

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Workflow - continued

- ▶ If yes, rearrange the asymptotic expansions to account for possible **nonlinear saturation** and get nonlinear amplitude equations (Landau-type for dispersive waves). Check saturation.
- ▶ Introduce (weak) **spatial modulation** and obtain Ginzburg-Landau (or nonlinear Schrödinger) type (dispersive waves), or Burgers-type (non-dispersive waves) evolution equations for wave amplitudes.
- ▶ Look for **pattern formation** in modulated amplitudes.

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Resonances leading to excitation of the trapped waves

- ▶ Free wave \rightarrow Trapped wave + Trapped wave
"Naive" perturbation theory \Rightarrow

$$A_{TT}^{(t)} = C |A^{(f)}|^2 A^{(t)}, \quad C > 0, \quad A^{(t)} \propto e^{cT}$$

- ▶ Free wave + Free wave \rightarrow Trapped wave
"Naive" perturbation theory \Rightarrow

$$A_T^{(t)} = CA_1^{(f)} A_2^{(f)}, \quad A^{(t)} \propto T$$

- ▶ Free wave + Mean flow \rightarrow Trapped wave
"Naive" perturbation theory \Rightarrow

$$A_T^{(t)} = CA^{(mf)} A^{(f)}, \quad A^{(t)} \propto T$$

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RSW in the presence of the coast

Non-dimensional equations of motion

$$\begin{aligned}u_t - v + h_x &= -\epsilon(uu_x + vv_y) \\v_t + u + h_y &= -\epsilon(uv_x + vv_y) \\h_t + u_x + v_y &= -\epsilon((hu)_x + (hv)_y).\end{aligned}\quad (12)$$

Boundary condition : $x \geq 0, \quad u|_{x=0} = 0.$

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Linear waves

Kelvin wave (KW)

$$(u, v, h) = (0, K(y + t), -K(y + t)) e^{-x}; \quad (13)$$

Inertia-gravity waves (IGW) :

Incident wave, $\sigma = \sqrt{1 + k^2 + l^2}$:

$$(u_i, v_i, h_i) = A_i \left(\frac{-k\sigma + il}{\sigma^2 - 1}, \frac{l\sigma + ik}{\sigma^2 - 1}, 1 \right) e^{i(l y - k x - \sigma t)} + \text{c.c.} \quad (14)$$

Reflected wave, $\sigma = \sqrt{1 + k^2 + l^2}$:

$$(u_r, v_r, h_r) = A_r \left(\frac{k\sigma + il}{\sigma^2 - 1}, \frac{l\sigma - ik}{\sigma^2 - 1}, 1 \right) e^{i(l y + k x - \sigma t)} + \text{c.c.}, \quad (15)$$

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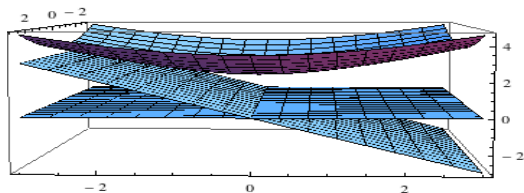
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Dispersion diagram



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IGW - KW resonance

A pair of IGW of the form is in resonance with a KW if

$$\sigma_1 - \sigma_2 = -l, \quad l_1 - l_2 = l, \quad l \neq 0. \quad (16)$$

We choose $l < 0$:

$$|l| = \sqrt{1 + k_1^2 + l_1^2} - \sqrt{1 + k_2^2 + l_2^2}, \quad l_2 = l_1 + |l|, \quad (17)$$

and

$$\sqrt{1 + k_1^2 + l_1^2} - |l| = \sqrt{1 + k_2^2 + (l_1 + |l|)^2}. \quad (18)$$

Algorithm for finding resonances :

1. Take any l and l_1 , then $l_2 = l_1 + |l|$,
2. Take arbitrary k_1 satisfying $k_1^2 \geq 2|l| \left(\sqrt{1 + l_2^2} + l_2 \right)$,
3. Define k_2 from

$$k_2^2 = k_1^2 - 2|l| \left(\sqrt{1 + k_1^2 + l_1^2} + l_1 \right).$$

Therefore, a KW with wavenumber l may be resonantly excited by a *continuum* of incident IGW with wavenumbers l_1 and

$|k_1| > \sqrt{2|l| \left(\sqrt{1 + (l_1 + |l|)^2} + l_1 + |l| \right)}$ interacting with another incident wave with k_2, l_2 :

$$k_2^2 = k_1^2 - 2|l| \left(\sqrt{1 + k_1^2 + l_1^2} + l_1 \right), \quad l_2 = l_1 + |l|. \quad (19)$$

Removal of resonances condition in RSW equations with coast :

$$\int_0^{\infty} dx e^{-x} (R_h - R_v) = 0, \quad (20)$$

where $R_{h,v}$ - r.h.s. of h - and v - equations.

Evolution equation for the amplitude of the Kelvin wave

$$K_T + KK_{\eta} = S e^{i\eta} + S^* e^{-i\eta}, \quad (21)$$

where $\eta = y + t$,

$$S = \int_0^{\infty} dx e^{-x} [(H_1 U_2^* + U_1 H_2^*)_x - U_1 V_{2x}^* - V_{1x} U_2 + i l (H_1 V_2^* + V_1 H_2^* - V_1 V_2^*)] \quad (22)$$

and (U_i, V_i, H_i) , $i = 1, 2$ are amplitudes of two IGW.

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Final form of the evolution equation

From the polarisation relations get :

$$S = iA_1 A_2 s, \quad lms = 0 \quad (23)$$

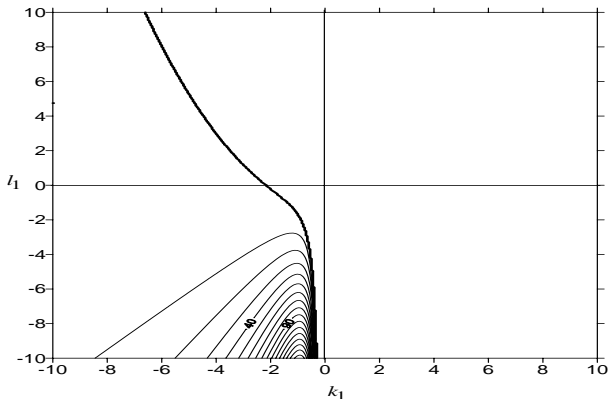
and hence

$$K_T + KK_{\eta} = -2sA_1 A_2 \sin l\eta, \quad (24)$$

where A_j - amplitudes of the two IGW,

$$s = \frac{4l}{(k_1^2 + 1)(k_2^2 + 1)[1 + (k_1 + k_2)^2][1 + (k_1 - k_2)^2]} \cdot \left[(\sigma_1 l_2 + \sigma_2 l_1 - l_1 l_2)(1 + k_1^2 + k_2^2) + \sigma_2 l_1 k_1(1 + k_1^2 - k_2^2) + \sigma_1 l_2 k_2(1 + k_2^2 - k_1^2) + 2k_1 k_2(l_1 l_2 - (1 + k_1^2)(1 + k_2^2)) \right] \quad (25)$$

Isopleths of the interaction coefficient $s(l, k_1, l_1)$ for $l = -1$ at the interval 10



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Integrability of the KW evolution equation

Forced simple-wave equation (after renormalizations) :

$$K_\tau + KK_\chi = -\sin \chi. \quad (26)$$

Lagrangian (characteristics) approach :

$$K = \mathcal{U} = \dot{\mathcal{X}}; \quad (\dot{\dots}) = \partial_\tau + \mathcal{U}\partial_\chi(\dots) \Rightarrow \quad (27)$$

$$\ddot{\mathcal{X}} + \sin \mathcal{X} = 0 \quad (28)$$

Pendulum equation : integrable. Shock formation \leftrightarrow

Lagrangian clustering (known in statistical physics :
mean-field limit of the kinetics of particles with repulsive
long-range interaction on the circle) \Rightarrow **Implications for
transport and mixing.**

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Ramifications :

- ▶ Including small dissipation \Rightarrow harmonically forced Burgers equation. Cole-Hopf change of variables \rightarrow Mathieu equation for Laplace transform in time \rightarrow integrable.
- ▶ Including small dispersion (long waves near a steep, but not vertical border) \Rightarrow harmonically forced KdV equation.

Summary

Resonant excitation of Kelvin waves by pairs of inertia-gravity waves near the coast is possible for a continuum of IGW - should be ubiquitous. The mechanism generates KW "from nothing". Subsequent slow evolution of KW leads to nontrivial transport and mixing properties.

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Equations of motion

2-layer rotating shallow water on the equatorial
tangent plane :

$$\partial_t \mathbf{u}_i + \mathbf{u}_i \cdot \nabla \mathbf{u}_i + \beta y \hat{\mathbf{z}} \times \mathbf{u}_i + \frac{1}{\rho_i} \nabla \pi_i = 0, i = 1, 2; \quad (29)$$

$$\partial_t h_i + \nabla \cdot (\mathbf{u}_i h_i) = 0, i = 1, 2, \quad (30)$$

$\mathbf{u}_i = (u_i(x, y, t), v_i(x, y, t))$ - velocity fields, h_i - depths of
the layers, $h_1 + h_2 = H$ - rigid lid upper b.c..

Barotropic and **baroclinic** velocities :

$$\mathbf{u}_{bt} = \frac{h_1 \mathbf{u}_1 + h_2 \mathbf{u}_2}{H}, \quad \mathbf{u}_{bc} = \mathbf{u}_1 - \mathbf{u}_2. \quad (31)$$

Parameters and characteristic scales

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$$q = \frac{H - 2H_1}{H}, \quad s = \frac{H_s}{H}, \quad \epsilon = \frac{\Delta H}{H_s}, \quad (32)$$

ΔH - typical variation of the interface, and $H_s = \frac{H_1(H-H_1)}{H}$.

Equatorial scaling :

$$L = \frac{(g'H_s)^{1/4}}{\sqrt{\beta}}; \quad T = \frac{1}{\beta L}; \quad U = \frac{g'\Delta H}{\beta L^2}. \quad (33)$$

Reduced gravity : $g' = g(\rho_2 - \rho_1)/\rho_1$

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Baroclinic/barotropic decomposition

$$\nabla^2 \psi_t + \psi_x = \epsilon \left[-J(\psi, \nabla^2 \psi) - \mathbf{s}(\partial_{xx} - \partial_{yy}) [(1 + \epsilon qh)(uv)] \right] \\ + \mathbf{s} \partial_{xy} \left[(1 + \epsilon qh)(u^2 - v^2) \right]$$

$$\mathbf{u}_t + \nabla h + \mathbf{y} \hat{\mathbf{z}} \times \mathbf{u} = \epsilon \left[-J(\psi, \mathbf{u}) + \mathbf{u} \cdot \nabla (\hat{\mathbf{z}} \times \nabla \psi) - q \mathbf{u} \cdot \nabla \mathbf{u} \right] \\ + \epsilon \mathbf{s} (2h \mathbf{u} \cdot \nabla \mathbf{u} + \mathbf{u} \mathbf{u} \cdot \nabla h),$$

$$h_t + \nabla \cdot \mathbf{u} = \epsilon \left[-J(\psi, h) - q \nabla \cdot (\mathbf{u} h) + \epsilon \mathbf{s} \nabla \cdot (h^2 \mathbf{u}) \right].$$

where ψ - barotropic streamfunction, $\mathbf{u} = (u, v)$ - baroclinic velocity and $h_1 = h$ - depth of the upper layer.

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Linear wave spectrum

- ▶ **Free barotropic** Rossby waves propagating at any angle :

$$\psi_0 = A_\psi e^{i(\theta+ly)} + c.c.; \quad \theta = kx - \sigma t; \quad (34)$$

with the dispersion relation

$$\sigma = -k/(k^2 + l^2). \quad (35)$$

- ▶ **Trapped baroclinic** waves

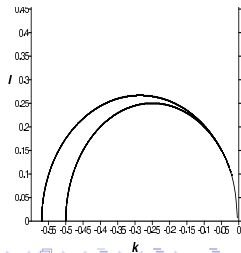
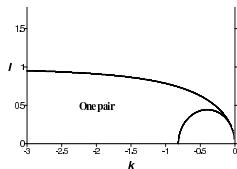
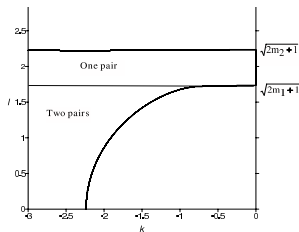
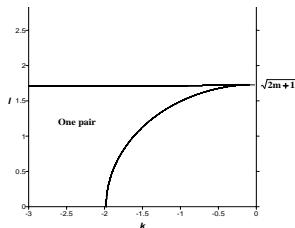
$$(u, v, h) = (iU_m, \phi_m, iH_m) A e^{i\theta_m} + c.c.; \quad \theta_m = \hat{k}x - \sigma_m t \quad (36)$$

with strongly localized at $y = 0$ amplitudes and dispersion relation

$$\hat{\sigma}_m^3 - (\hat{k}^2 + 2m + 1)\hat{\sigma}_m - \hat{k} = 0; \quad m = 0, 1, 2, \dots \quad (37)$$

Solutions to the synchronism conditions

Lecture 4: Resonant interactions and resonant excitation of waveguide modes



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First nonlinear correction to the triad :

$$\begin{aligned}\nabla^2 \psi_t^{(1)} + \psi_x^{(1)} &= -\nabla^2 \psi_T^{(0)} + N_\psi, \\ u_t^{(1)} - yv^{(1)} + h_x^{(1)} &= -u_T^{(0)} + N_u, \\ v_t^{(1)} + yu^{(1)} + h_y^{(1)} &= -v_T^{(0)} + N_v, \\ h_t^{(1)} + u_x^{(1)} + v_y^{(1)} &= -h_T^{(0)} + N_h,\end{aligned}$$

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Nonlinear source-terms

$$N_\psi = -J(\psi^{(0)}, \nabla^2 \psi^{(0)}) - \mathbf{s}(\partial_{xx} - \partial_{yy})(u^{(0)} v^{(0)}) \\ + \mathbf{s} \partial_{xy} (u^{(0)2} - v^{(0)2}),$$

$$N_u = \frac{-J(\psi^{(0)}, u^{(0)}) + \mathbf{u}^{(0)} \cdot \nabla \psi_y^{(0)}}{} - q \mathbf{u}^{(0)} \cdot \nabla u^{(0)}$$

$$N_v = \frac{-J(\psi^{(0)}, v^{(0)}) - \mathbf{u}^{(0)} \cdot \nabla \psi_x^{(0)}}{} - q \mathbf{u}^{(0)} \cdot \nabla v^{(0)}$$

$$N_h = \frac{-J(\psi^{(0)}, h^{(0)})}{\underline{\hspace{10em}}} - q \nabla \cdot (\mathbf{u}^{(0)} h^{(0)}).$$

Only underlined terms may contain resonances.

Removal of resonances :

$$A_{1T} = L_1 A_\psi \bar{A}_2, \quad \bar{A}_{2T} = \bar{L}_2 \bar{A}_\psi A_1,$$

These equations are reduced to a single one :

$$A_{\alpha TT} = C |A_\psi|^2 A_\alpha, \quad C = L_1 \bar{L}_2.$$

C is real. An important property is $C > 0 \Rightarrow$ **exponential growth.**

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Energy balance

$$E = E_{bt} + E_{bc} = \text{const},$$

where E_{bt} , E_{bc} are the barotropic and the baroclinic energies :

$$E_{bt} = \int_{-\infty}^{\infty} dy \langle (\nabla \psi)^2 \rangle_x, \quad E_{bc} = \frac{S}{2} \int_{-\infty}^{\infty} dy \langle (1 + \epsilon qh)(u^2 + v^2 + h^2) \rangle_x.$$

Quadratic energy form conserved in the lowest order :

$$E_0 = \frac{S}{2} \int_{-\infty}^{\infty} dy \langle (u^{(0)2} + v^{(0)2} + h^{(0)2}) \rangle_x + \int_{-\infty}^{\infty} dy \langle \nabla \psi^{(0)} \cdot \nabla \psi^{(1)} \rangle_x,$$

First barotropic correction (the barotropic response of the equator) is crucial

"Pure" parametric resonance $\sigma = 2\hat{\sigma}$, $k = 2\hat{k}$.

Amplitude (Landau) equation :

$$A_{T_2} + LA_\psi \bar{A} + (P + iQ) |A|^2 A = 0.$$

$Q = Q_0 + qQ_1 + sQ_2$, L, P, Q are real, $P \geq 0$, $T_2 = \epsilon^2 t$.

Term $\propto L$: the interaction of the primary barotropic wave with the zero-order baroclinic one,

Term $\propto P + iQ_0$: interaction between the secondary barotropic mode and the zero-order baroclinic mode,

Term with Q_1 : interaction between zero- and first-order baroclinic fields,

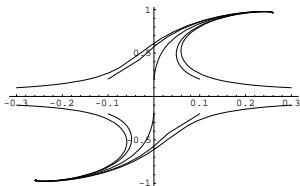
Term with Q_2 : cubic interaction of the zero-order baroclinic mode.

Stationary solutions :

$$A_0 = 0, A_{\pm}^2 = -\frac{LA_{\psi}}{P + iQ}.$$

A_0 is unstable, A_{\pm} are **stable and attractive** if $P > 0$ (neutrally stable if $P = 0$). \Rightarrow **nonlinear saturation always takes place.**

Typical behaviour of the trajectories of solutions of the Landau equation in the phase-space $ReA - ImA$



Two different baroclinic waves.

$$A_{1T_2} + \alpha_1 \bar{A}_2 + \beta_1 |A_2|^2 A_1 + \gamma_1 |A_1|^2 A_1 = 0$$

$$A_{2T_2} + \alpha_2 \bar{A}_1 + \beta_2 |A_1|^2 A_2 + \gamma_2 |A_2|^2 A_2 = 0.$$

The following properties of the coefficients may be established :

$$\alpha_1 \bar{\alpha}_2 > 0, \operatorname{Re}(\beta_1 + \beta_2) \geq 0, \operatorname{Re}(\gamma_1) \geq 0, \operatorname{Re}(\gamma_2) \geq 0.$$

$\alpha_{1,2}$ are either both real or both imaginary, and at least one of $\operatorname{Re}(\gamma_1), \operatorname{Re}(\gamma_2)$ is > 0 . \Rightarrow saturation.

Solution - **attracting limit cycle** :

$$A_{1,2} = B_{1,2} e^{\pm i\omega T_2}$$

"Pure" parametric resonance $\sigma = 2\hat{\sigma}$, $k = 2\hat{k}$ with spatial modulation

Amplitude equation in the reference frame moving with the trapped wave :

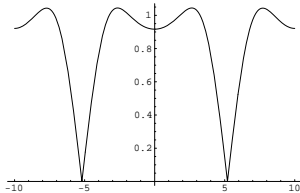
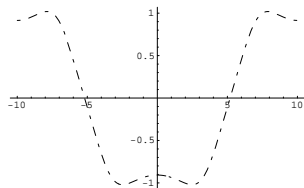
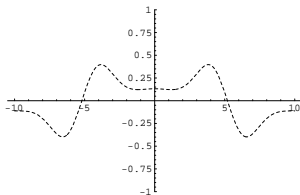
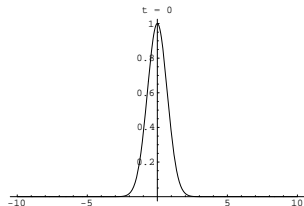
$$A_{T_2} - \frac{i}{2} \hat{\sigma}''(\hat{k}) A_{X_1 X_1} - LA_{\psi} \bar{A} + (P + iQ) |A|^2 A = 0.$$

$\hat{\sigma}''(\hat{k})$ is the derivative of the group velocity with respect to \hat{k} , $X_1 = \epsilon X$.

Equation of the **Ginzburg - Landau** (GL) type : resonantly forced GL equations known in optics and in the Faraday effect.

The stationary solutions are still solutions, stable for $P > 0$. Two different stationary states - characteristic Bloch-type **domain wall** structures appear, forming a bound state (a so-called bubble, or **"dark soliton"**)

Domain walls and dark soliton



Initial (top left), and final profiles of real (dashed) and imaginary (dash-dotted) parts of A , and $|A|$ (solid)

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A pair of baroclinic waves

The amplitude equations in the lowest order :

$$A_{i\tau_1} + c_{g_i} A_{ix_1} = 0, \quad i = 1, 2,$$

c_{g_i} - group velocities of the respective waves. Next order :

$$\begin{aligned} A_{1\tau_2} + c_{g_1} A_{1x_2} - \frac{i}{2} \sigma_1''(k_1) A_{1x_1x_1} + \alpha_1 \bar{A}_2 + \beta_1 |A_2|^2 A_1 \\ + \gamma_1 |A_1|^2 A_1 = 0 \\ A_{2\tau_2} + c_{g_2} A_{2x_2} - \frac{i}{2} \sigma_2''(k_2) A_{2x_1x_1} + \alpha_2 \bar{A}_1 + \beta_2 |A_1|^2 A_2 \\ + \gamma_2 |A_2|^2 A_2 = 0, \end{aligned}$$

$$X_2 = \epsilon^2 X.$$

Comments

- ▶ Resonant excitation of trapped waves by non-trapped ones in *semitransparent* waveguide = parametric instability of the barotropic wave with respect to baroclinic perturbations.
- ▶ New type of GL equation : spatio-temporal organization ("dark solitons")
- ▶ Coupled GL-type equations : slow variability, fronts.

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Summary of the results

What we have seen :

- ▶ Incident IGW can generate coastal Kelvin waves "from nothing"
- ▶ Barotropic planetary waves passing through equator excite equatorial Rossby and Yanai waves

What we have not seen, but can be shown :

- ▶ Shelf and edge waves can be generated by IGW similarly to Kelvin waves
- ▶ Generation of trapped waves both in coastal and equatorial wave-guides can be provided by interaction of a free wave with a mean current, although resonance conditions are more restrictive.

Discussion

Importance

- ▶ Teleconnections midlatitudes-tropics
- ▶ Slow variability in the tropical atmosphere
- ▶ Transport of energy in the ocean from the large to the coasts \Rightarrow a route for dissipation
- ▶ Transport and mixing properties in the coastal zones

Realisability

Weakly nonlinear phenomenon : efficiency conditioned by persistent of free-wave flux towards the wave-guide.

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